# Overview of plasma diagnostics (high energy density plasmas)

S.V. Lebedev

**Imperial College London** 



### **General comments**



HEDP systems *emit*:

- Visible, XUV, x-ray photons
- Charged particles
- Neutrons

Can be *probed* by similar particles

Small objects – high spatial and temporal resolution (10 $\mu$ m, 1ns)

Principles of plasma diagnostics are mostly common for all types of plasmas



**Electrical measurements: magnetic field and current** 

Laser probing: interferometry, density gradients

**Emission: spectral lines, continuum** 

**Thomson scattering** 

X-ray imaging

**Proton probing** 

### **Magnetic field**





#### **Magnetic coil**

$$\dot{\vec{E}}_{C} \vec{E} \cdot d\vec{l} = -\iint_{S} \dot{\vec{B}} \cdot d\vec{s} \qquad V = NA\dot{B}$$

after integrator:

$$B(t) = \frac{RC}{NA} \cdot V_{\text{int}}(t)$$

Example of magnetic probe designed to measure magnetic field distribution inside the plasma

#### More often a set of coils outside

Always check absence of capacitive coupling! (rotation of coil by 180<sup>o</sup> should give identical signal but of opposite polarity)





### **Magnetic field**



**Magnetic coil** 

$$\dot{\vec{E}}_{C} \vec{E} \cdot d\vec{l} = -\iint_{S} \dot{\vec{B}} \cdot d\vec{s} \qquad V = NA\dot{B}$$

after integrator:

 $B(t) = \frac{RC}{NA} \cdot V_{\text{int}}(t)$ 

Faraday

rotation

Farada

Capacitor-coi target

a = 250 µm

Probe beam

9 ns

TNSA proton beam

up to 20 MeV

Magnetic probes in HEDP:

- Usually positioned at some distance from the object - require extrapolation of the measured signal to the object (sometimes by many orders of magnitude)
- Can be strongly affected by a "noise" from energetic particles generated in the system (can be tested using two oppositely-wound probes)



Measured: 5 mT at 3cm from the coil Extrapolated: 800 T in the coil  $\rightarrow$  800/5x10<sup>-3</sup> = 1.6x10<sup>5</sup>



**B**-probe

Proton-

RCF

Wollaston

**Driver laser beam** 

1ns, square time-profile 500 I. 1017 W/cm2

Streak Camera

stack

deflectometry

**B-dot** 

probe

### Rogowski coil



Gives total current through the loop, independent on current distribution.

(if the signal is not "too fast", i.e.  $\tau$  > propagation time along the coil)

$$\Phi = n \oint_l \int_A dAB \cdot dl \qquad \qquad \oint_l B \cdot dl = \mu_0 I$$

$$\Phi = nA\mu_0 I \qquad \qquad V = \dot{\Phi} = nA\mu_0$$

- n is number of turns per unit length, A is area of the coil cross-section
- Induced voltage is proportional to dl/dt (V can be integrated before recording)
- "Return wire" to exclude contribution from magnetic flux through the coil
- needs electrostatic shielding





**Rotation of polarisation plane (Faraday effect)** 



#### **Propagation of electro-magnetic waves**

 $V_{ph} = \frac{\omega}{k} = \frac{c}{\eta}$   $\eta \neq 1 \Rightarrow$  phase shift (interferometry)  $\nabla \eta \neq 0 \Rightarrow$  refraction of the beam (schlieren and sadowgraphy)  $\eta_{+} \neq \eta_{-} \Rightarrow$  rotation of polarisation plane (Faraday effect)

**Refractive index of plasma without magnetic field:** 

$$\omega^{2} = \omega_{p}^{2} + k^{2}c^{2} \implies \eta = \sqrt{\left(1 - \frac{\omega_{p}^{2}}{\omega^{2}}\right)} = \sqrt{\left(1 - \frac{n_{e}}{n_{cr}}\right)} \qquad n_{cr}\left[cm^{-3}\right] = \frac{1.12 \cdot 10^{21}}{\lambda^{2}\left[\mu m\right]}$$

Probing beam cannot propagate if the plasma density is above the critical density Cut-off densities:

- $\lambda=337\mu m$  (HCN laser)  $~n_{cr}$  =  $10^{16} cm^{-3}$
- $\lambda=$  10.6  $\mu m$  (CO\_2 laser)  $n_{cr}$  = 10^{19} cm^{-3}
- $\lambda = 1.06 \mu m$  (Nd,  $\omega_0$ )  $n_{cr} = 10^{21} cm^{-3}$
- $\lambda=0.532\mu m$  (Nd  $2\omega_0) ~n_{cr}$  = 4  $^{\cdot}10^{21} cm^{-3}$

### Interferometry









In

Phase shift (typically n<sub>e</sub><<n<sub>cr</sub>):

$$\varphi = \int (k_p dl - k_r dl)$$
$$\varphi = \frac{\omega}{c} \int \left[ \sqrt{1 - \frac{n_e}{n_{cr}}} - 1 \right] dl \approx -\frac{\omega}{2cn_{cr}} \int n_e dl$$

Number of fringes:

$$F = \frac{\varphi}{2\pi} = 4.46 \cdot 10^{-18} \lambda_{[\mu m]} \cdot \int_{[cm^{-2}]} n_e dl$$

#### **Michelson configurations**

1.5 -

1.0 -

0.5 -

0.0 -



Line density corresponding to one fringe:

- $\lambda = 337 \mu m$  (HCN laser)  $\langle n_e L \rangle = 6.6 \cdot 10^{14} cm^{-2}$
- $\lambda = 10.6 \mu m (CO_2 \text{ laser}) \quad \langle n_e L \rangle = 2.1 \cdot 10^{16} \text{ cm}^{-2}$

$$\lambda = 1.06 \mu m \text{ (Nd, } \omega_0 \text{)} \qquad \langle n_e L \rangle = 2.1 \cdot 10^{17} \text{ cm}^{-2}$$

 $\lambda = 0.532 \mu m (Nd 2\omega_0)$   $\langle n_e L \rangle = 4.2 \cdot 10^{17} cm^{-2}$ 

For "cw" plasmas need to exclude vibrations (two wavelength interferometry)

### Interferometry



#### **Refractive index for neutral gases**

$$\eta_a = 1 + (2\pi e^2 / m) \cdot \sum_{i,k} \frac{f_{i,k} n_i}{(\omega_{ik}^2 - \omega^2)}$$

# For optical frequencies far from spectral lines ( $\alpha$ is polarizibility):

$$\eta_a \approx 1 + 2\pi\alpha \cdot n_a$$

Atom	<b>2</b> πα
Не	1.3 <sup>.</sup> 10 <sup>-24</sup> cm <sup>3</sup>
H <sub>2</sub>	5 <sup>.</sup> 10 <sup>-24</sup> cm <sup>3</sup>
Air	1.1 <sup>.</sup> 10 <sup>-23</sup> cm <sup>3</sup>
Ar	1.10 <sup>-23</sup> cm <sup>3</sup>
ΑΙ	4.4 <sup>.</sup> 10 <sup>-23</sup> cm <sup>3</sup>

 $f_{i,k}$  is an oscillator strength for i to k transition  $\omega_{ik}$  is the frequency of transition  $n_i$  is the density of atoms in state i

#### **Phase shift:**

$$\varphi = \frac{2\pi}{\lambda} \int (1 - \eta) dl$$

Number of fringes due to electrons and atoms:

$$F = \frac{\varphi}{2\pi} = -4.46 \cdot 10^{-14} \cdot \lambda_{[cm]} \cdot \int n_e dl + \frac{2\pi\alpha}{\lambda} \cdot \int n_a dl$$

•different sign of phase shift for electrons and atoms [e.g. for H<sub>2</sub>  $\phi$ =0 at n<sub>e</sub>/n<sub>a</sub>=2.2% (0.56% for He) @ 6943Å] •different dependence on  $\lambda$ 

### Interferometry







Phase shift (typically n<sub>e</sub><<n<sub>cr</sub>):

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$$\varphi = \frac{\omega}{c} \int \left[ \sqrt{1 - \frac{n_e}{n_{cr}}} - 1 \right] dl \approx -\frac{\omega}{2cn_{cr}} \int n_e dl$$

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For "cw" plasmas need to exclude vibrations (two wavelength interferometry)

### Interferometric imaging







 $\downarrow \\ \mbox{reference fringes ( d = \lambda / [2 \cdots in(\alpha/2)] ) in the regions without plasma} \\ \mbox{Fringe shift due to plasma} \Rightarrow map of \langle n_e L \rangle \\ \mbox{Practical lower limit $$\sim$1/10 - 1/30 of fringe} \end{cases}$ 

Misalignment between the beams ( $\alpha$ )







s.lebedev@ic.ac.uk May 2019

a particular position

Numbers show fringe-shift in

## Simple Example: Interferogram of 8 Wire Tungsten Array



# Analysis method: [Swadling et al. - PoP 20, 022705 (2013)] Trace experimental and background fringes:



# Number fringes, interpolate and subtract



# More Complex example – 32 wire Al array

Results published - *Swadling et al.* - Physics of Plasmas 20, 022705 (2013).



# More Complex example – 32 wire Al array



### Schlieren and shadow imaging





**Deflection angle:** 

$$\alpha = \frac{d}{dy} \int \eta dl \approx \frac{1}{2n_{cr}} \frac{d}{dy} \int n_e dl$$

Example:

α=2·10<sup>-3</sup>rad, λ=532nm ∫⊽n<sub>e</sub>dl = 1.5·10<sup>19</sup>cm<sup>-3</sup> Knife to measure 1-D gradients pin-hole for "bright-field" schlieren disk for "dark-field" schlieren



#### Intensity distribution

 $\frac{\Delta I}{I} = L \left[ \frac{d^2}{dx^2} + \frac{d^2}{dy^2} \right] \int \eta dl$ 

(If an imaging system between the object and the detection plane is used, then *L* is the plasma length)

### **Example: plasma jet – obstacle interaction**



Supersonic magnetized plasma jet interacts with an obstacle (experiments on MAGPIE facility at Imperial College):

 $n_{e} \sim 10^{18} cm^{-3}$ ,  $V_{jet} \sim 100 km/s$ ,

 $\alpha$  = 5.10<sup>-3</sup>rad,  $\lambda$ =532nm

Schlieren image highlights the positions of sharp gradients: bow-shock and collimation shock at the jet boundary

Interferogram allows measurements of the electron density in the regions where interference fringes are traceable

### **Faraday rotation**





Linearly polarised wave is superposition of two circularly polarised.

Difference in refractive indexes of the two waves leads to rotation of polarisation plane

#### **Rotation of polarisation plane:**

$$\alpha = \frac{e}{2m_e c} \int \frac{n_e B \cdot dl}{n_{cr} (1 - n_e / n_{cr})^{1/2}} \approx \frac{e}{2m_e c} \int \frac{n_e}{n_{cr}} \vec{B} \cdot d\vec{l}$$

$$\alpha_{[rad]} = 2.26 \cdot 10^{-17} \lambda_{[cm]}^2 \int n_{e[cm^{-3}]} \vec{B}_{[G]} \cdot d\vec{l}$$

Example:

 $n_e$ =10<sup>19</sup> cm<sup>-3</sup>, B=10<sup>5</sup>G (10T),  $\lambda$ =532nm, *L*=1cm =>  $\alpha$  =3.67<sup>o</sup>

# Magnetic field can be found if the density distribution is known

#### Convenient to combine with interferometry



M. Tatarakis et al, PoP, 5, 682 (1998)

### **Faraday rotation**





### **Example of Faraday rotation measurements**

### Interaction with conducting obstacle



### Polarimetry images: two channels



### **Abel inversion**





Many diagnostics measure a particular plasma parameter integrated along probing path.

For a cylindrically symmetric object it is possible to reconstruct radial distribution from several chord integrals

$$F(y) = \int f(r) dx = 2 \int_{y}^{a} f(r) \frac{r dr}{\sqrt{(r^{2} - y^{2})}}$$

Chord integral of e.g. phase shift, rotation angle or radiative emissivity

$$f(r) = -\frac{1}{\pi} \int_{r}^{a} \frac{dF}{dy} \frac{dy}{\sqrt{(y^2 - r^2)}}$$

Radial profile of quantity F

Accuracy of reconstruction strongly depends on errors in dF/dy  $\Rightarrow$  need to have sufficiently large number of chords and low level of "noise"



**Electrical measurements: magnetic field and current** 

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X-ray imaging

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### **Continuum radiation**



Free-free (bremsstrahlung)

Free-bound (recombination on level n)

 $j_{ff}(v) \sim n_e n_i Z_i^2 T^{-0.5} e^{-hv/kT_e}$ 

 $j_{fb}(v) \sim n_e n_i Z_i^4 T^{-3/2} e^{(I_n - hv)/kT_e}$ 

j(v) is spectral power per unit frequency divide by  $\lambda^2$  if need  $j(\lambda)$ !

Temperature measurements from the slop of continuum radiation ( $h_V > kT_e$  – typically from X-ray spectrum)



Density ( $Z_{eff}$ ) measurements from absolute intensity (for  $h_V << kT_e$  – in visible ) if temperature is known

(weak dependence on temperature)

$$j(\lambda) \sim \frac{n_e^2 Z_{eff}}{T^{1/2}} \qquad \qquad \sum_i n_e n_i Z_i^2 \equiv n_e^2 Z_{eff}$$

need to chose spectral region free of lines



### **Continuum radiation**



#### **Black-body spectrum**

$$j_{bb}(v) \sim \frac{hv^3}{e^{hv/kT} - 1}$$

$$j(\lambda) = \frac{1.19 \cdot 10^{20}}{\lambda_{A}^{5} \cdot (e^{\left(\frac{12395}{\lambda_{A}T_{ev}}\right)} - 1)} \left[ W / cm^{2} / \dot{A} / ster \right]$$

**Transmission Grating Spectrometer** 



#### Radiation spectrum of tungsten Z-pinch Cuneo et al, PoP 2001



### **Relative intensities of spectral lines**



$$j = \frac{h v_{ki}}{4\pi} A_{ki} n_k$$

$$A_{ki} = 6.67 \cdot 10^{15} \frac{g_i}{g_k} \frac{f_{ik}}{\lambda_{\dot{A}}^2}$$

A<sub>ki</sub> is transition probability for spontaneous emission [s<sup>-1</sup>]

 $n_k$  is number of atoms (ions) in the upper state [cm<sup>-3</sup>]

 $\boldsymbol{f}_{ik}$  is the "oscillator strength", usually used in spectroscopy

Need to know how population of different states depends on plasma parameters!

#### Local Thermodynamic Equilibrium (LTE).

$$\frac{n_k}{n_0} = \frac{g_k}{g_0} \exp(-\frac{E_k}{kT})$$

$$\Downarrow$$

**Boltzmann distribution** 

$$\frac{j_1}{j_2} = \frac{A_1}{A_2} \frac{g_1}{g_2} \exp(-\frac{\Delta E}{kT})$$

Temperature from the ratio of line intensities

$$n_e[cm^{-3}] > 9 \cdot 10^{17} \left(\frac{T}{E_H}\right)^{1/2} \left(\frac{\Delta E}{E_H}\right)^3$$
  
 $(E_H = 13.6eV)$ 

LTE is a good approximation at high density, when collisional transitions dominate over radiative

Griem, H.R., "Plasma spectroscopy"

### **Relative intensities of spectral lines**

#### **Coronal Equilibrium**

$$\frac{dn_k}{dt} = n_e n_0 \langle \sigma_{1k} V \rangle - \sum_{j \neq k} n_k A_{kj} = 0$$

Collisional excitation balanced by radiative de-excitation

$$\frac{n_k}{n_0} = \frac{n_e \langle \sigma_{0k} V \rangle}{\sum_j A_{kj}} \propto f(T)$$

Population of levels depends on temperature

**Collisional Radiative models** (equilibrium or time-dependent)







### X-ray spectrum of a 600eV AI plasma





Effect of density on emission spectrum from 3mm diameter Al plasma

Lines and continuum

•Saturation at blackbody level

•Line "inversion" at high densities due to self-absorption

## X-ray spectrum of a 600eV AI plasma









1000 photon energy (eV) 10000

### **Spectral line broadening**



#### **Doppler broadening**

Line profile for Maxwellian distribution:

$$I(v) = I(v_0) \exp\left(-\frac{(v - v_0)^2 c^2}{2V_T^2 v_0^2}\right)$$

Also for measurements of directed velocity Problem - choice of line (in X-rays  $\Delta\lambda$  is too small) Ion temperature from  $\Delta\lambda/\lambda$ 

$$T_i[eV] = 1.68 \cdot 10^8 \frac{m_i}{m_H} \left(\frac{\Delta \lambda_{1/2}}{\lambda_0}\right)^2$$

e.g. for 
$$H_{\beta}$$
 at  $T_i$ =100eV  $\Delta\lambda \sim 3.7$ Å

#### **Stark broadening (in electric field produced by neighbouring particles)**

**Characteristic E field:** 

Characteristic  $\Delta\lambda$  (FWHM) (for H<sub> $\beta$ </sub> ( $\lambda$ =4861 Å)): e.g.  $\Delta\lambda \sim 1.8$  Å for n<sub>i</sub>=10<sup>15</sup>cm<sup>-3</sup>

$$E \propto \frac{e}{r^2} \sim n_i^{2/3}$$

$$\Delta \lambda_{1/2} [\dot{A}] = 0.4 \cdot \left( \frac{n[cm^{-3}]}{10^{14}} \right)^{2/3}$$



### **Zeeman splitting**



$$\Delta \lambda_{[A]} = (M_2 g_2 - M_1 g_1) \lambda_{[A]}^2 \cdot 4.7 \cdot 10^{-13} B_{[G]}$$

Example: Li beam excited by dye laser



Usually difficult to find a suitable spectral line!

(  $\Delta\lambda/\lambda \sim \lambda$  – hard in X-rays )

### **Zeeman splitting**



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$$\Delta \lambda_{[A]} = (M_2 g_2 - M_1 g_1) \lambda_{[A]}^2 \cdot 4.7 \cdot 10^{-13} B_{[G]}$$

#### Example: laser plasma in magnetic field

Al III 4s-4p ( ${}^{2}S_{1/2} - {}^{2}P_{1/2}$  vs  ${}^{2}S_{1/2} - {}^{2}P_{3/2}$ ) Different sensitivity to B-field: Zeeman splitting – B-field. Stark broadening – n<sub>e</sub>





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 $|k_s| \approx |k_i|$ 

laser

### Thomson scattering (optical)

Scattering of e/m radiation by single electron

#### scattering geometry

k<sub>i</sub>, k<sub>s</sub> are wave-vectors of incident and scattered radiation V is velocity vector of an electron

 $\frac{d\sigma}{d\Omega} = r_e^2 \sin^2(\phi)$ 

 $\sigma = \frac{8\pi}{2} r_e^2 = 6.65 \cdot 10^{-25} cm^2$ 



velocity || to vector k





Scattering by plasma



Total electric field in the scattered wave is the sum of electric fields scattered from a large number of small plasma cells (each with density  $n_e=n_e+\delta n_e$  and phase shift  $\varphi_i$ )

$$E_{s} \sim \sum_{j} (\langle n_{e} \rangle + \delta n_{e}) exp(-i\phi_{j}) \sim \sum_{j} \delta n_{e} exp(-i\phi_{j}) \qquad \left( \langle n_{e} \rangle \sum_{j} exp(-i\phi_{j}) = 0 \right)$$

 $I_{s} \sim \langle E_{s} E_{s}^{*} \rangle \sim \sum_{jk} \langle \delta n_{k} \delta n_{j} \rangle \exp\left(-i(\phi_{k} - \phi_{j})\right) \implies$ 

Scattering signal is determined by correlation of density fluctuations in plasma

Plasma quasi-neutrality (Debye shielding):

- **D** <  $\lambda_{D}$  density fluctuations from thermal motion
- $D > \lambda_D$  from collective modes (waves)

### **Scattering parameter**





 $\alpha <<1$  ( $\Lambda << \lambda_D$ ) => scattering on fluctuations occurring on spatial scales  $<< \lambda_D$ , thermal fluctuations inside Debye sphere (incoherent scattering)

 $\alpha >>1$  ( $\Lambda >> \lambda_D$ ) => scattering on collective modes, from electrons which motions are correlated ( coherent scattering )



#### Thomson scattering spectral density function

$$\sigma = \sigma_T \cdot S(\vec{k}, \omega) \qquad (\omega \equiv \omega_s - \omega_i)$$
$$S(\vec{k}, \omega) = S_e(\vec{k}, \omega) + S_i(\vec{k}, \omega)$$



# [Froula *et al.* "Plasma scattering of electromagnetic radiation", 2010]

#### **Qualitatively:**

"lon" and "electron" terms in the total scattering -EM are scattered by the electrons (ions are too heavy)!

#### Debye shielding in plasma

Electron motions are followed by "electron cloud" (+e): Fast motions, large frequency shifts, "electron" term

Ion motions shielded by electron cloud (-e/2) and by "ion cloud" (-e/2)

Slow motions, small frequency shifts, "ion" term





$$S_{e}(k,\omega) = \int f(\vec{V}) \delta(\omega - \vec{k}\vec{V}) d\vec{V}$$
$$S_{e} = n_{e} \sqrt{\frac{m_{e}}{2\pi T_{e}k^{2}}} \exp\left(-\frac{m_{e}\omega^{2}}{2T_{e}k^{2}}\right)$$
$$\frac{\Delta\lambda_{1/2}}{\lambda} = 6.6 \cdot 10^{-3} \sqrt{T_{e[eV]}} \sin(\theta/2)$$

<u>Incoherent scattering</u>: fluctuations are determined by thermal motion of electrons (low density, high temperature, large scattering angles)

- f(V) is the electron distribution function
- for Maxwellian distribution

e.g.,  $\Delta\lambda$  = 25nm for Nd laser (532nm) at 90<sup>o</sup> (for plasma with T<sub>e</sub>=100eV)





Possibility to measure electron distribution function (e.g. presence of fast electrons)







Total scattering (frequency integrated)



#### **Coherent scattering:**

scattering from test electron is balanced by scattering from the shielding (electron) cloud  $\Rightarrow$  S<sub>e</sub> => 0

for a test ion scattering is from the shielding electron cloud  $\Rightarrow$  S<sub>i</sub> dominates ("scattering on ions") (high density, low temperature, small scattering angles)



Narrow ion feature centred on broad electron spectrum. Electron peaks give plasma density:

$$(\Delta \omega)^2 = \omega_p^2 (1 + 3/\alpha^2)$$



Ion wings due to ion acoustic resonance

$$\Delta \omega = k \left( \frac{\alpha^2}{1 + \alpha^2} \frac{k_B T_e}{m_i} + \frac{3k_B T_i}{m_i} \right)^{1/2}$$
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s.lebedev@ic.ac.uk May 2019



**Experimental considerations** 



Amplitude of the signal (lower limit on density)

$$\frac{P_s}{P_{Laser}} \sim n_e \sigma_T L \Delta \Omega \quad \sim 10^{-10} for n_e \sim 10^{18} cm^{-3}$$

Plasma light (upper limit on density)

$$P_{S} \propto n_{e}$$
  $P_{ff,fb} \propto n_{e}^{2}$ 

Scattering geometry – plasma light is collected from a larger volume than scattered



Need to suppress scattering from the walls, windows and optics

or separate by time-of-flight for short laser pulses



**Experimental considerations** 



Spatially resolved scattering using fibreoptics and imaging spectrometer

Measurements of plasma flow velocity and plasma temperature

(MAGPIE facility at Imperial College)

[Harvey-Thompson et al., PRL 2012 & PoP 2012]





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X-ray imaging

**Proton probing** 

## X-ray imaging



#### Pin-hole camera





Imploding Z-pinch



**Detector:** 

X-ray film (time-integrated) or MCP camera (time-resolved)

**Detector irradiation:** 

$$\frac{energy}{unit area} \propto \frac{\Delta^2}{q^2}$$

### X-ray radiography



**Point-projection** 





Detector: film (time-integrated) or MCP



### X-ray spectroscopy and radiography with spherical crystals



#### **Spectroscopy with 1D spatial resolution**







Spectrum of AI Z-pinch on Magpie



### **Probing with ion beams**





**Deflection of ions by electric or magnetic fields** For small angles: ٢

$$\sin(\theta) = \frac{q}{\sqrt{2m_i E_i}} \cdot \int B \times dl \approx 7 \cdot \frac{\int B \times dl [T \cdot m]}{\sqrt{A \cdot E[MeV]}}$$

$$\tan(\theta) = \frac{q}{2E_i} \cdot \int E_\perp dl \approx \frac{1}{2E_i[MeV]} \cdot \int E_\perp dl \ [MV / m \cdot m]$$

Ionisation



MCF: probing by heavy ions (A=200, E~300keV) Measurements of electric potential in plasma: position of ionisation from trajectory potential from change of energy

### **Proton probing**





Proton beam from laser irradiated foil (~MeV with broad energy spectrum)

Registration on stack of radiohromic films, to provide energy resolution

- Point-projection imaging with MeV energy protons
- Proton attenuation is typically negligible, deflection by E- and B-fields
- Broad energy spectrum of protons.
- Different layers in the film stack correspond to different proton energy can be used for time-resolved imaging (ps time-of-flight delay)

### **Proton probing**



Macchi et al., RMP, v.85 (2013)



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- **Proton attenuation is typically negligible, deflection by E- and B-fields** ٠
- Broad energy spectrum of protons. ٠
- Different layers in the film stack correspond to different proton energy can ٠ be used for time-resolved imaging (ps time-of-flight delay)

## **Proton probing**





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